# Warming of early Mars induced by CO<sub>2</sub> ice clouds:

Estimations of cloud condensation flux, column density and radius by a one-dimensional radiation model

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# 1. Introduction: The faint young sun paradox on Mars

### **Present: cool and dry climate**

Major component : CO<sub>2</sub>
Atmospheric pressure : 0.006 atm
Surface temperature : 216 K
liquid water not to exist stability



Valley networks

### 38 Gyr ago: warm and wet climate

Major component : CO<sub>2</sub> (photochemical stability)

Valley networks exist: Denser atmosphere and higher surface temperature enough for liquid water to exist stably

# Young Sun was dark. "The faint young sun paradox"

Earth: Explainable if CO<sub>2</sub> pressure had been higher (Kuhn and Kasting, 1983)

Mars: Unexplainable because upper limit exists in atompspheric pressure. (Kasting, 1991)

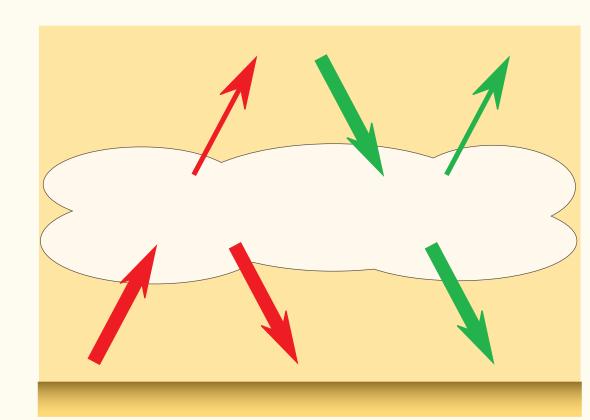
- 1-D radiative convective model ( $CO_2$   $H_2O$  atmosphere)
- Consider vertical temperature construction changed by atmospheric condensation.

  (But he neglected the radition processes of clouds)
- The surface freazing and collapse condensation couse

if atmospheric pressure more than the upper limit.

# 2. The scattering greenhouse effect of CO<sub>2</sub> ice clouds

Pierrehumbert and Erlick (1998)



If the backward scattering of the planetary radiation is larger than that of solar radiation by the clouds, we can expect climate warming.

# Scattering greenhouse effect of cloud

Cloud particle radius : 10  $\sim$  20  $\mu$ m (they can effectively reflect IR radiation ) Climate warming strongly

Previous studies (Mischna et al. 2002; Yokohata et al. 2002; Colaperete and Toon, 2003)

- · The level of the greenhouse effect strongly depends on cloud parameters
  - (particle radius and column density)
- · Climate become warm for appropriate values of cloud parameters

### However

- · The particle radius dependency of surface temperature has not been examined
- · Feasibility of the such values has not been examined

# Goal of this study:

Investigation of radius dependency on surface temperature Estimations of the column density and particle radius

### 3. The vertical construction and 1-D radiative transfer model

Atmospheric components: CO<sub>2</sub>, H<sub>2</sub>O

Radiation equilibrium

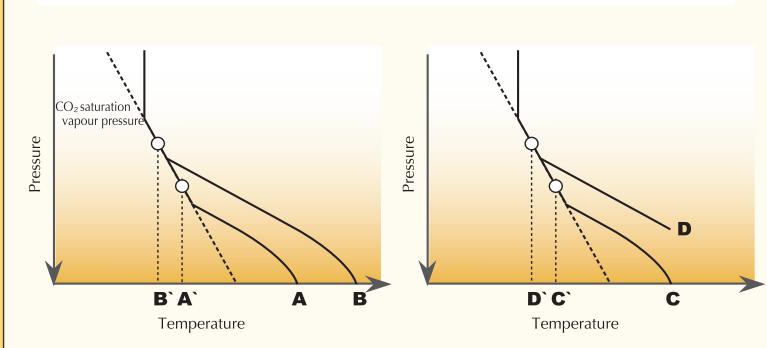
(CO<sub>2</sub> saturation vapour pressure curve)

(Thin grey layer)

CO<sub>2</sub> MOIST ADIABAT

Solar luminosity: 75 % as present

Surface albedo: 0.216



H<sub>2</sub>O MOIST ADIABAT

Temperature

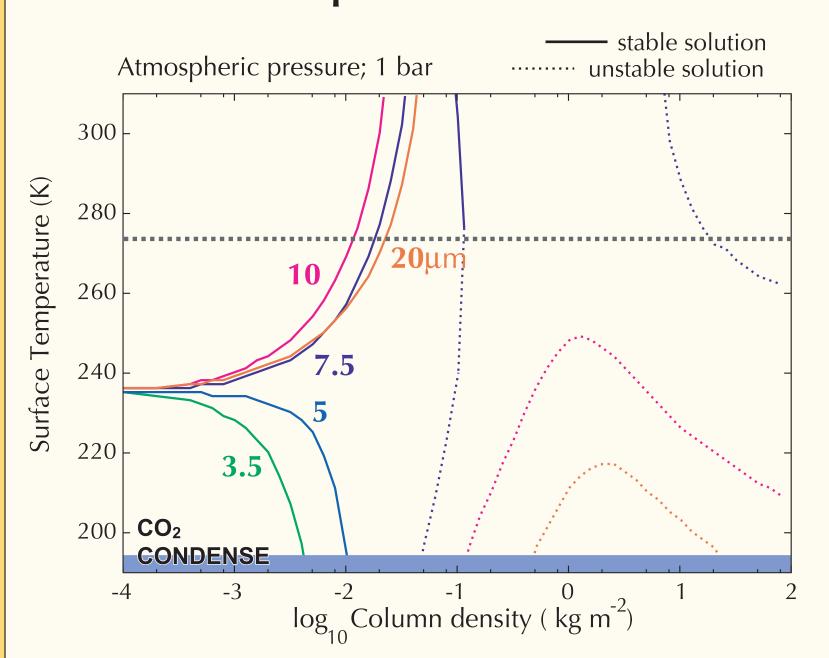
Stratosphere:

Cloud layer: Delta-edington approximation
Cloud particle: Mie theory (assuming spherical particles)
Complex indices of CO<sub>2</sub> ice (Warren, 1986)
Gas: the random model
band parameters (Houghton, 2002)

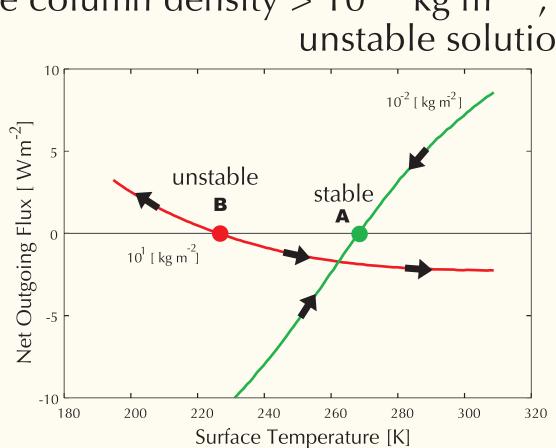
band parameters (Houghton, 2002)
Gas-only layer: Two-stream approximation
Line-by-line method (CO<sub>2</sub>, H<sub>2</sub>O)
absorption line parameters (HITRAN2000)

### 4. Results and Discussion

#### 4.1 Surface temperature under radiative balance - the particle radius dependency -



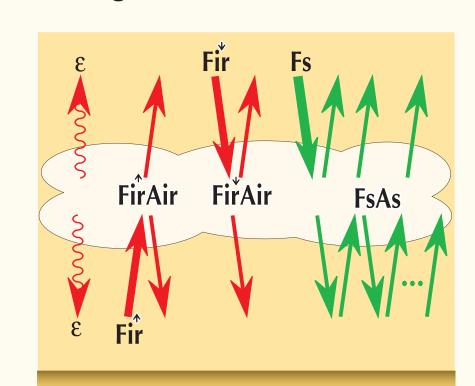
- 5 μm: anti-greenhouse effect 7.5 μm - 20 μm: greenhouse effect, the column density  $> 10^{-1}~kg~m^{-2}$ ; unstable solution

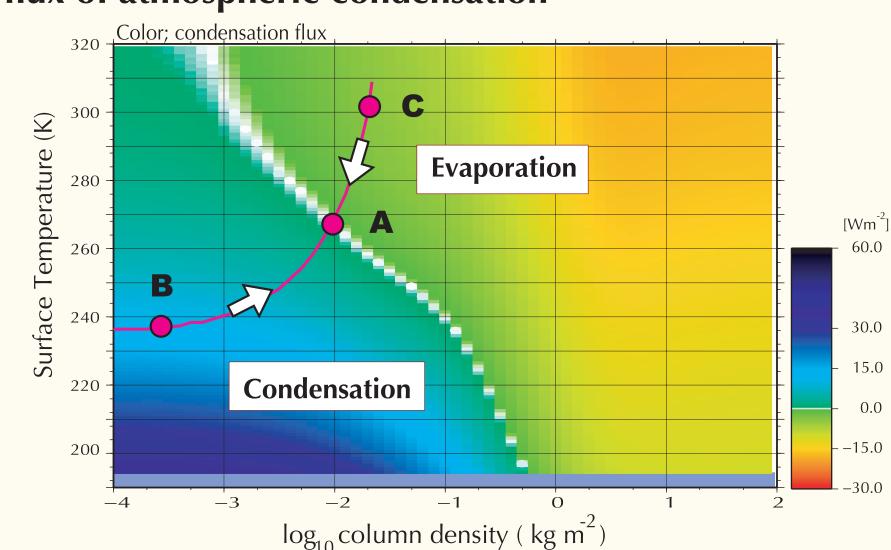


#### 4.2 Condensation flux; energy flux of atmospheric condensation

Net radiative cooling energy in cloud layer ||
Condensation flux

 $F_L = 2\varepsilon - (Fir^+ + Fir^-) Air - FsAs$ 





Condensation flux decreases

when surface temperature and column density increase

- Negative feedback mechanism of codensation flux
- for cloud column density change makes climate stable
- condensasion-evaporation equilibrium is archived
   (But, neglect particle radius changed by condensation)
- Under this equilibrium the cloud column density and particle radius have a relationship (they were treated independent parameters!)

#### 4.3 Estimations of the column density and particle radius by condensation flux

log<sub>10</sub> Column density (kg m<sup>-2</sup>)

Column number density  $10^9 \sim 10^{10} \,\mathrm{m}^{-2}$ ; Surface temperature  $\sim 270 \,\mathrm{K}$ 

4.4 The column density decreases by evaporation as getting out of the cloud

Estimation by Yokohata et al. (2002)

- Neglect vertical distribution of density
- · Particles go down by Stokes settling velocity

5 10 20μm 20μm consider consider log<sub>10</sub> Column density [ kg m<sup>2</sup>]

This effect poor influences on estimations of cloud parameters

### 5. Conclusion

getting out of the cloud.

We estimate the level of the greenhouse effect and the cloud parameters by the 1-D radiation model under  $CO_2$ -H<sub>2</sub>O atmosphere which is assumed as early Martian atmosphere.

Negative feedback mechanism of codensation flux for changing in cloud column density makes climate stable.

- We can estimate the column density as the function of the particle radius or the column number density

- When atmospheric pressure is fixed as 1 bar,

- + Column density  $1.0 \times 10^{-2} \text{ kg m}^{-2}$ : (@ radius  $10 \text{ }\mu\text{m}$ )
- surface temperature 268 K
- + Column density 1.0 x 10 kg m<sup>-2</sup>: (@ column number density  $10^{10}$ m<sup>-2</sup>) surface temperature 270 K These estimations do not changed when we consider the column density decreases by evaporation as

The minimum requisites to induce the warm and wet climate:

Atmospheric pressure is more than about 1 bar, the column number density nearby equals to  $10^9 \sim 10^{10} \text{m}^{-2}$